

#### CMOS Solid-State Photomultipliers for High Energy Resolution Calorimeters

#### **Optical Detector with Integrated ADC for Digital Readout**

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# **PRIMEX Calorimeter Readout**

- Utilizing the Jefferson Lab Upgrade.
- Provide direct measurements at low energies of parameters of Quantum Chromodynamics (QCD) using η and η' lifetime measurements.
- Primakoff effect production of neutral mesons, which decay.
- Need to measure position and energy deposition in calorimeter.





- About 1% energy resolution at 4.5 GeV
- Replace lead glass with smaller PbWO<sub>4</sub>
- Cost effective readout will make this possible.



# **The PRIMEX PbWO<sub>4</sub> Calorimeter**

- Planned Calorimeter
  - ♦ 60 x 60 element array of PbWO<sub>4</sub>
  - <1% energy resolution for 4.5 GeV</p>
  - ♦ ~ 1 mm position resolution
  - ◆ 2.125 x 2.125 x 21.5 cm<sup>3</sup>
- Detecting two high energy gamma rays
  - Scattering along scintillator
  - Scattering laterally
  - Bundle clusters of scintillators





# **Building the Calorimeter**



# **Solid-State Photomultipliers**



- Solid-state photomultiplier (SSPM) is an array of Geiger photodiodes read out in parallel.
- Each photodiode has a gain of 10<sup>6</sup> for single photon events.
- Number of diodes triggered is proportional to the incident light flash.
- Compact and phototube like response.





# **Single Geiger Photodiode Response**



- Breakdown: ~26.7 V
- Range of Operation: 10 V above breakdown.
- $\succ C_{jn} = 150 \, pF$
- > Gain @  $V_x = 1$  V: 9x10<sup>5</sup>
- *τ*<sub>rise</sub> ≈ 1.7 ns
- *τ*<sub>fall</sub> ≈ 30 ns

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## **Detection Efficiency**

- Detection efficiency is a product of the QE and the Geiger probability.
- Difference in ionization rates between holes and electrons.
- > There may be differences in the Geiger avalanche probability,  $P_g$ , as a function of wavelength.
- Many scintillation materials emit in the blue.
- Small changes in the DE for blue light can result in a significant improvement in the signal.



### **Excess Noise Analysis**



- Conducted an extensive analysis on processing the waveforms.
- Robust method for characterizing each noise term.
- Identify differences between expected and measure dark spectra.
- Distinguish dark noise, cross talk, and after pulsing.

# **Geiger-Mode Multiplier**

$$\boldsymbol{n}_{t}(\tau) = \boldsymbol{M}_{X} \cdot \boldsymbol{M}_{AP}(\tau) \cdot \boldsymbol{n}_{d}(\tau) = \boldsymbol{M} \cdot \boldsymbol{n}_{d}(\tau)$$

Spectral Mean

$$n_{d}(\tau) = n_{ttl} \left( 1 - \left[ \frac{N_{0}}{N_{T}} \right]^{\frac{1}{n_{ttl}}} \right)$$
 Dark Triggers  
Binomial

- ➢ All terms are averages.
- Estimate total number of possible diode triggers based on integration time and number of diodes in the SSPM.
- Determine spectral mean.
- Calculate expected mean.
- Ratio gives the multiplier.
- Determine excess noise factor, but we are still looking at it.

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- Short Integration:
  - $M_{\chi}$  is given
  - Scales only with bias
- Long Integration:
  - Product of  $M_X$  and  $M_{AP}$

# **SSPM Connection and Amplifier**





#### SSPM Function:

- Evaluated each component
   of the modular readout boards.
- SSPM connections were tested.
- Simple trans-impedance amp and a light pulse from an LED was used.

- Amplifier Function:
  - Each photodetector on the device is connected properly. The fast amplifier used is a 2.2 GHz differential amplifier.
  - The signal from the SSPM is readout after the amplifier.



### **Boost Converter**

- > A DC/DC boost converter is used to power the SSPM.
- The voltage will be controlled through an interface through the FPGA.
- Each circuit will need to be calibrated.
- Filtering is used to remove the switching noise from the converter.



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22 uH

Vbb

MBR0540





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#### **Excess Bias Monitor**



- A large isolated pixel is used to monitor the excess bias.
- The voltage drop across the diode is held by a slow RC circuit.



- The dark current is sufficiently large to keep the voltage stable.
- The response is linear will applied bias over a 5 V swing.
- The offset is accurate to within 2.5 mV.



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# **Complete Functionality**







- Coupled to an FPGA evaluation board.
- Placed a LYSO crystal on the SSPM and powered the device through the boost converter.
- The 12-bit word from the ADC was sent into the FPGA and processed.
- The code calculated an integral for each gamma event from a 22Na source.
- The integrals were saved for a file and a histogram was generated, reproducing the expected results.

#### **Determining the Expected Performance**

We want to provide a model that will accurately predict the behavior of an SSPM-based detector.



- Excess noise is the unknown factor.
- Readout Scheme:
  - Large gain from the SSPM allows for a very simple readout COTS components.
  - Fast electronic components are necessary for extracting the physics information.
  - Evaluation components were built and tested at 0 °C.



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# **Data Capture for Characterization**



# LaBr<sub>3</sub> on PMT

- No readily available high-energy gamma ray source for evaluation with PbWO<sub>4</sub>.
- Used a high-quality, fast, bright scintillator.
  - Lanthanum Bromide.
  - ♦ ~60 photons/keV
  - Decay Time: 16 ns
- Evaluate with PMT
  - Optically coupled with grease.
  - Crystal in aluminum can.
  - Super bialkili cathode
  - ◆ QE<sub>Eff</sub>: ~ 32.8%
- Scintillator Contribution:
  - ◆ 511 keV: 6.1% (FWHM)
  - ◆ 1275 keV: 4.4% (FWHM)

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# **SSPM Spectra**

- Generated spectrum as it will be done in the PRIMEX experiment.
- Collect wave forms and process.
- ➤ Use a 200 ns integration window.
- Find maximum.
  - Capture a waveform over 8 μs.
  - Step over waveform with 200 ns window.
  - Find point with largest integration value.
  - Generate histogram of maximum values.
- Measurements:
  - Temperature: 0 °C
  - Bias: 29, 30, 31, 32, 33 V
  - Optically coupled using grease.
  - Exposed with gammas from <sup>22</sup>Na.

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Counts

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## **Excess Noise Factor**

- > Know quantities:
  - PDE and DCR from SSPM
  - Scintillator response
- Scaling factor, F, is used to account for excess noise:
  - After pulsing
  - Cross talk
- Systematic Error:
  - Light yield
  - PMT response
  - Varied assumptions and used average.
- At an excess bias of 6 V (P<sub>g</sub> ~60%), the excess noise factor is 1.26.
- > F of 2 at  $V_x = 7.5 \text{ V}, P_g \sim 70\%$



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# Lead Tungstate

- CMS experiment at CERN
  - Lead Tungstate based calorimeter (ECAL)
  - Use Hamamatsu APDs
  - Extract APD expected performance from measured resolution.
- PRIMEX upgrade
  - Replace lead glass.
  - Higher density of scintillators.
  - To be operated at 0 C.
- Lead Tungstate
  - Decay Time: ~ 16 ns
  - Light yield: 151 photons/MeV at 0 C.
  - Peak emission: 480 nm





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# **Calorimeter Performance**

- Calorimeter geometry:
  - 2.05 cm x 2.05 cm crystals
  - Use 4, 0.8 cm x 0.8 cm SSPM.
  - Accounting for packaging
- Physics:
  - High energy gamma ray
  - Electron cloud: Moliere Radius
  - 5 x 5 scintillator array is used for each gamma ray interaction
- SSPM performance:
  - ♦ Include PDE, DCR, F
  - Time response: pixels may trigger more than once
  - Scintillator light yield and decay
- Calorimeter performance:
  - Fold in scintillator response based on CMS measurements
  - Scintillator is the limiting factor
  - Speculation at this point: Needs to be measured in high energy gamma field



### **FPGA Board**

- Two prototype FPGA evaluation board has been fabricated.
  - Mounts four SSPM modules.
  - ADC with MUX for digitizing excess bias monitoring signals.
  - ◆ 1 USB port for data transfer.
- The board has not shorts to ground.
- The parts have been purchased and assembly has begun.
- The assembly will be done next month and evaluation will be immediately on each component.



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# Summary

- The specified SSPMs are being designed into our next submission.
- We are looking at the optical interfaces of the SSPMs to improve the photon collection.
- The SSPM module works; minor design revisions will be implemented to reduce cost and improve performance.
- Digitizing the signal at the detector will significantly reduce cost while preserving sufficient signal to noise for the experiment.
- > This work is expanding the capabilities for RMD:
  - Serving the DOE and other government agencies.
  - Developing instruments for broader applications.



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